

PHY 127

Prof. Ben Kilminster

Lecture 2

Mar. 1st, 2024

Did you already take PHY117 (or the equivalent of two semesters of univers

Value	Count	%
Yes	144	81%
No	34	19%

How many semesters of physics did you take in Gymnasium ?

Value	Count	%
0	9	5%
1	2	1%
2	51	29%
3+	116	65%

What is your major ?

Value	Count	%
biomedicine	169	95%
biology	7	4%
biodiversity	0	0%
some type of chemistry	0	0%
some type of social science	1	1%
other	1	1%

What faculty is your minor ?

Value	Count	%
no minor or faculty of science	154	87%
faculty of theology	2	1%
faculty of law	1	1%
faculty of medicine	9	5%
faculty of business, economics, informatics faculty of veterinary medicine	12	7%

Specifically, what is your minor ?

Minors:

- astronomy/astrobiology (6)
- neuroinformatics (6)
- biology (3)
- business (3)
- bioinformatics (3)
- computer science
- geography (2)
- ethics (2)
- economics (2)
- Biomedicine (2)
- chemistry
- law
- physics
- statistics

What do you want to learn ?

Much of what you asked for is already in this lecture, so we are all set. Here are some common themes:

“physics that matters in our daily lives and how things work”

“understand human body and universe”

“quantum physics that doesn't happen in our daily lives”

“increase knowledge of how things work”

“practical use of physics in scientific research and innovation”

“how medical diagnostics work”

Here are a few things mentioned, and I will try to explain them along the way:

“how to measure how far something is away from earth/the speed of it – generally speaking galaxy stuff”

“string theory”

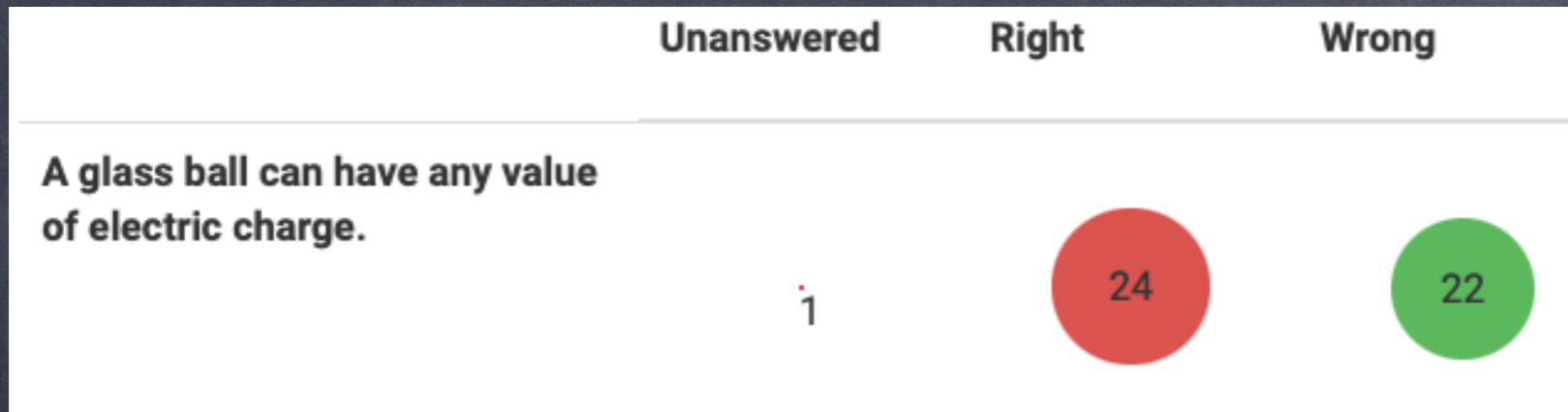
“soft-body matter, condensed matter”

“LHC”

“DNA sequencing”

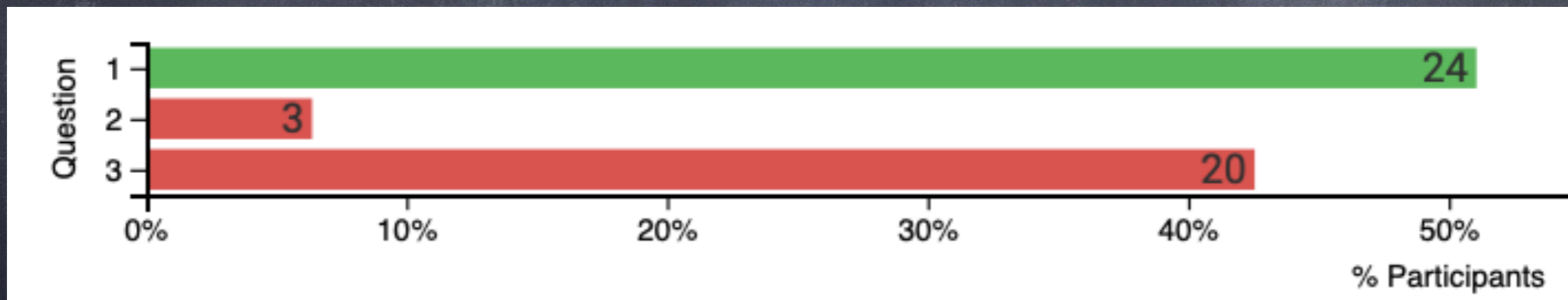
“how a camera works”

and of course: “how to pass exam” (15)

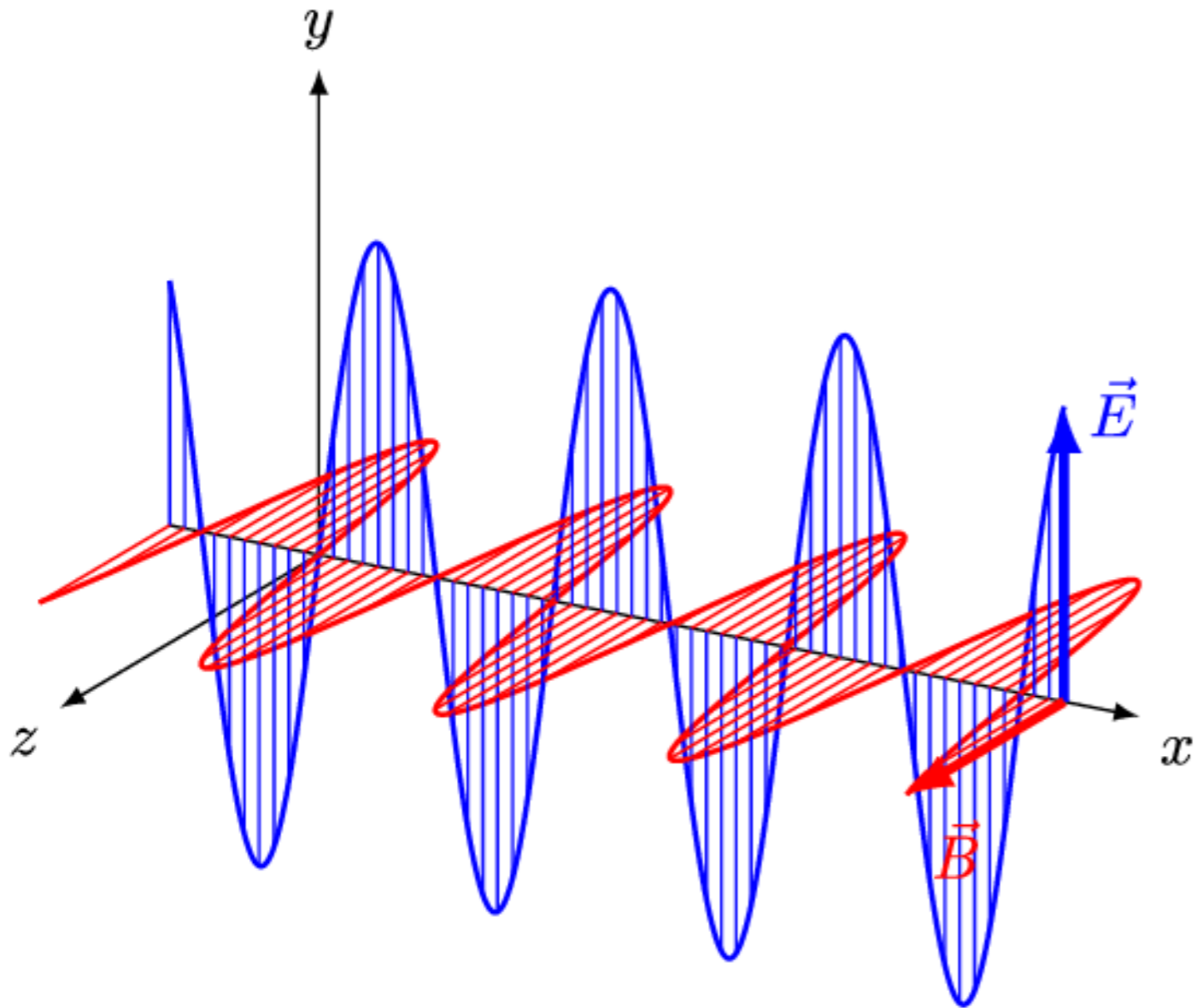


Practically speaking, which of the above options would yield the most precise value of the radius ?

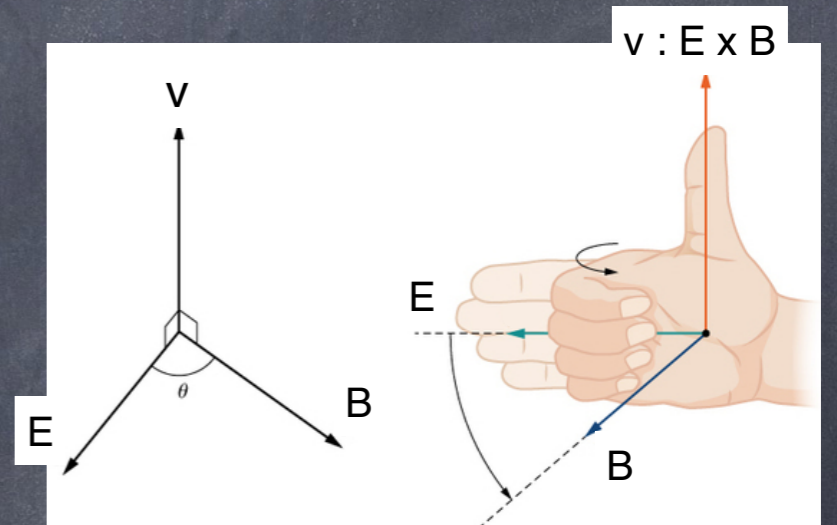
- We measure how fast it falls in a swimming pool and use this to calculate the ball's size.
- We measure how much higher the water is in the swimming pool with and without the ball and then determine the radius of the ball.
- We put a specific amount of electric charge on the ball, then put two charged metal plates on the top and bottom of the swimming pool, and then change the voltage on the plates until the ball is suspended.



At end of PH4 17, you learned light is an electromagnetic wave. The amplitude of \vec{E} is in \hat{y} -direction. The magnetic field is \perp to \vec{E} field, & is in \hat{z} -direction. The direction of propagation of light is $\vec{V} \sim \vec{E} \times \vec{B}$

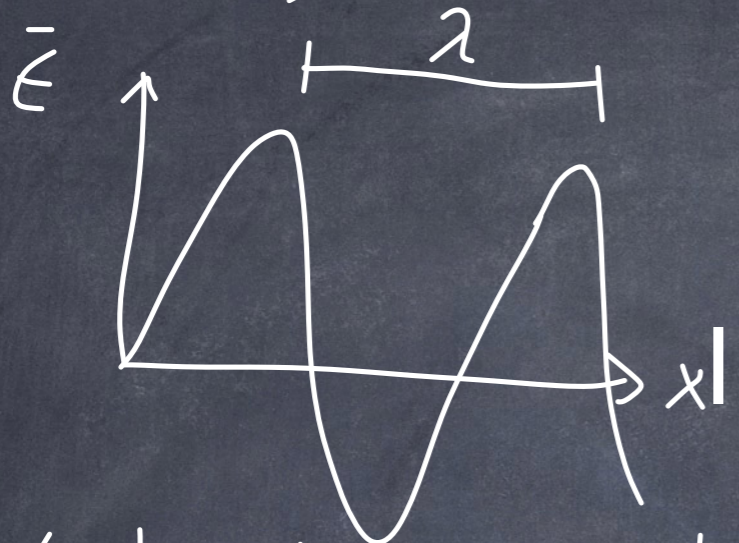


\vec{V} is in direction of \hat{x} -direction



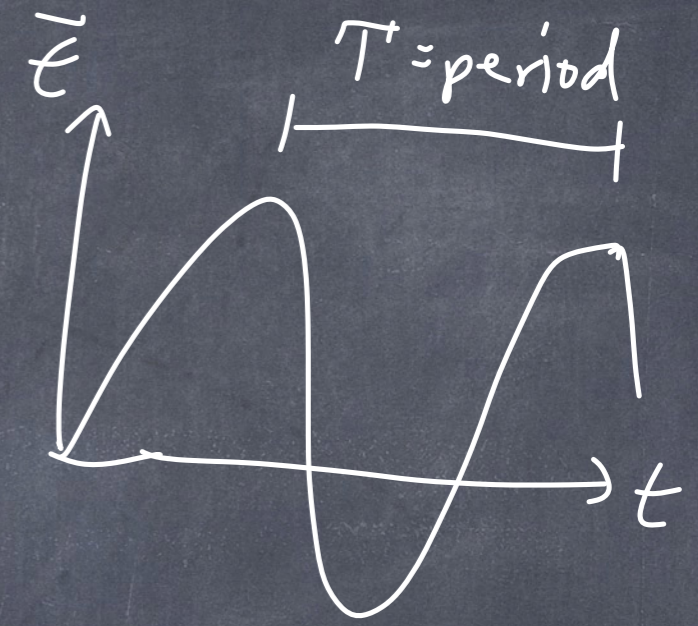
Review on cross products & unit vectors is in script physics 1, Chapter 3.

Focussing on \vec{E} , it changes with position and time.



Light has a wavelength, λ
units of $[m]$

and has a frequency, ν
units of $[\frac{1}{s}]$
Hz

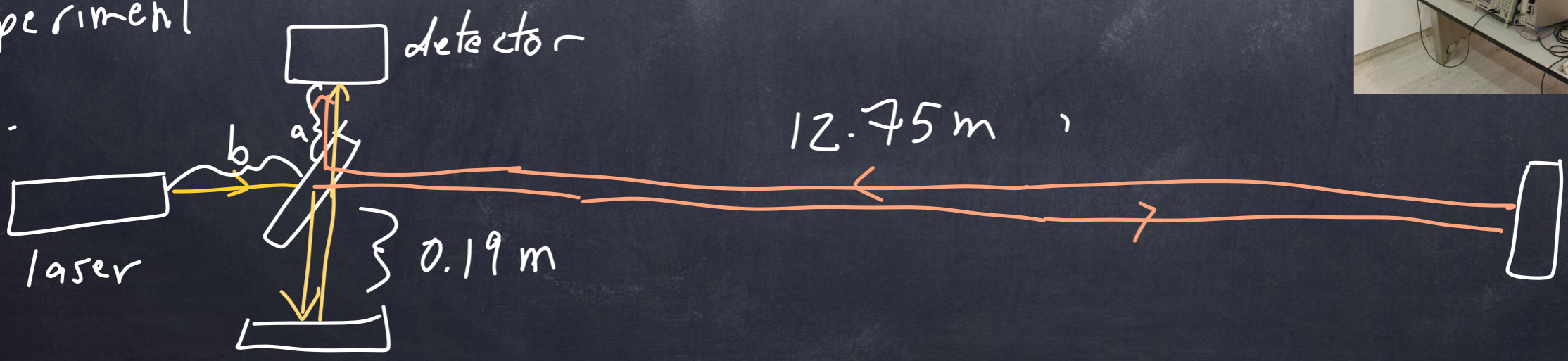


$$\nu = \frac{1}{T}$$

For a wave, $v = \lambda \nu$ $[\frac{m}{s}]$
velocity

For light,

Experiment



distance of path 1: $b + 2(12.75\text{ m}) + a$
(longer)

distance of path 2: $b + 2(0.19\text{ m}) + a$

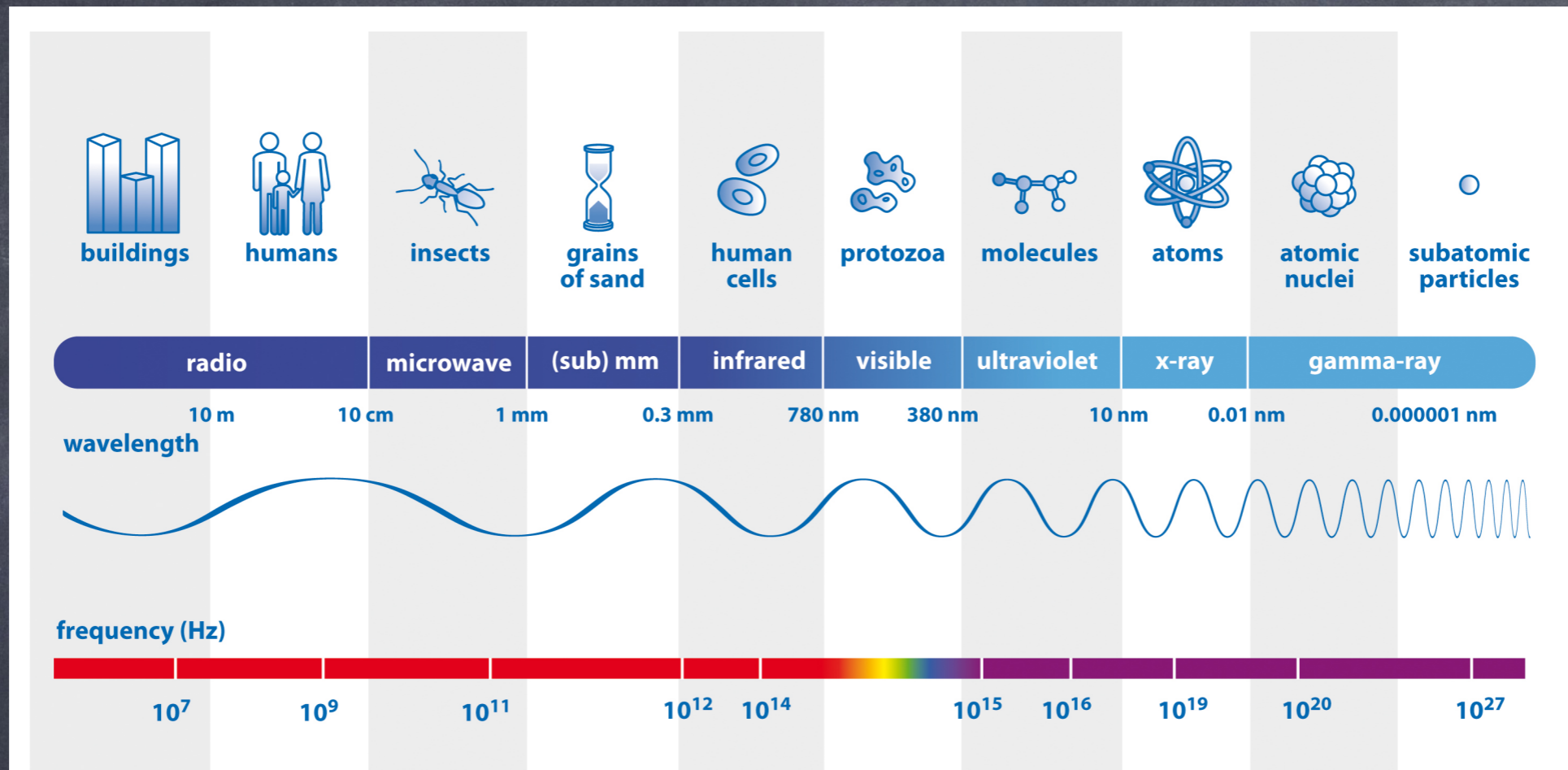
$$\Delta x = \text{path 1} - \text{path 2} = \cancel{b} + \cancel{a} + 2(12.75\text{ m}) - (\cancel{b} + \cancel{a} + 2(0.19\text{ m}))$$
$$= 2(12.75\text{ m}) - 2(0.19\text{ m}) = 25.12\text{ m}$$

$$\Delta t = \text{measured} = 84 \text{ E-9 s}$$

$$c = \frac{\Delta x}{\Delta t} = \frac{25.12\text{ m}}{84 \text{ E-9 s}} = 2.99 \text{ E 8 } \frac{\text{m}}{\text{s}}$$

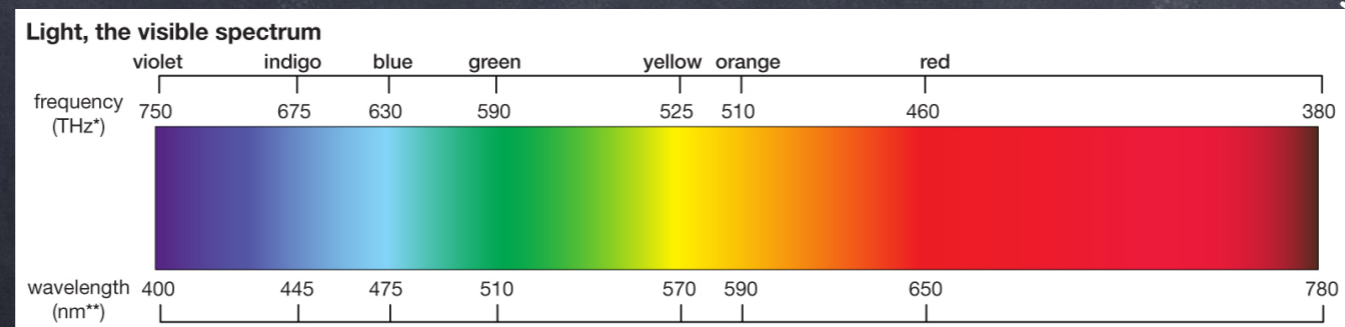
$$\downarrow$$
$$2.998 \dots \frac{\text{m}}{\text{s}}$$

Light can have many wavelengths + frequencies.



↓ (flipped)

ultraviolet



infrared

↑
400 nm

↑
700 nm

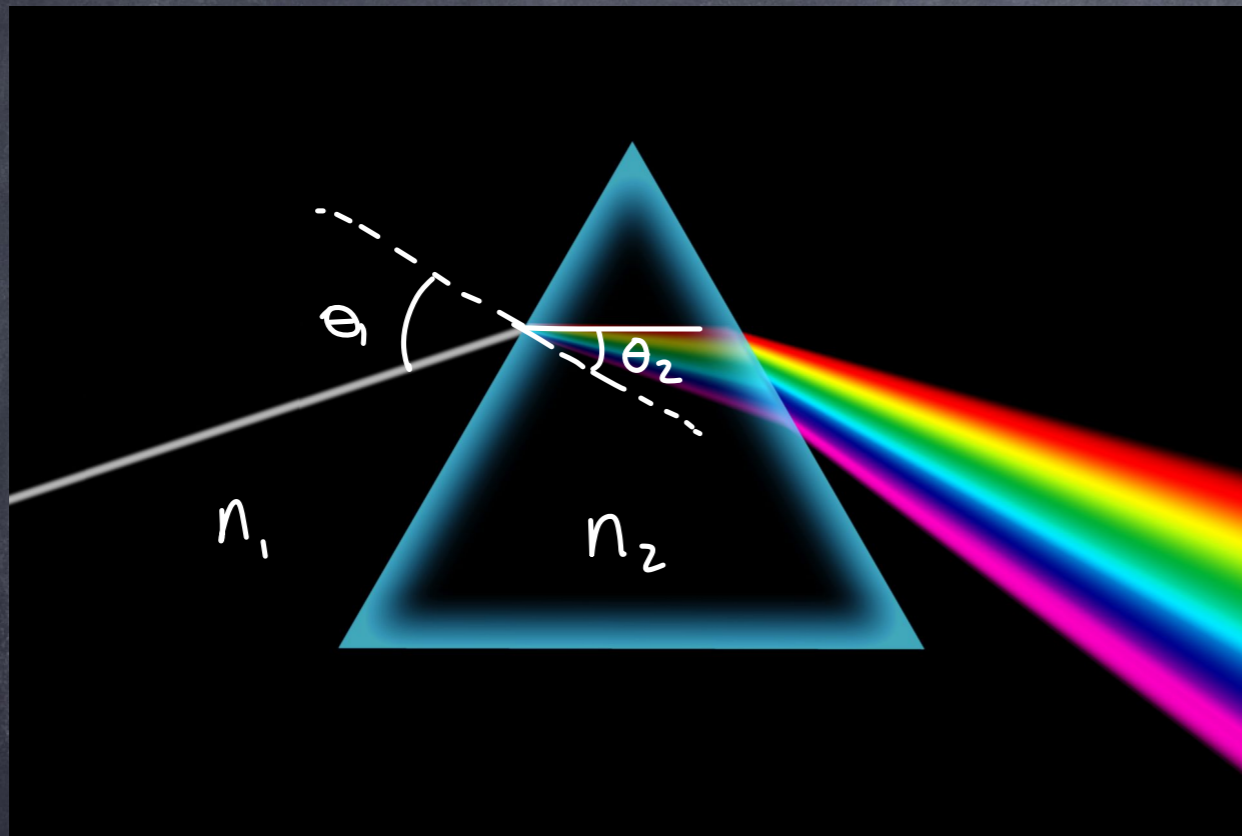
Experiment: split light

PHY 117:

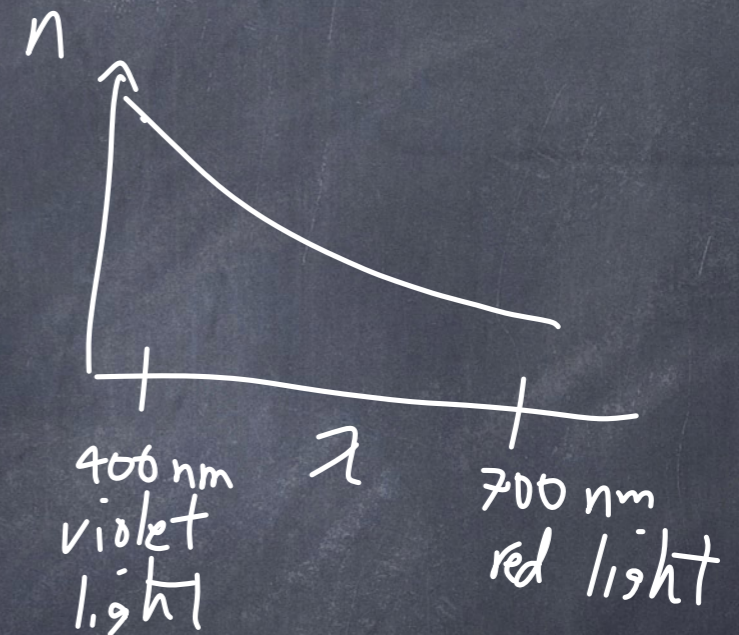
Snell's Law:

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$

n : index of refraction



n depends on the wavelength of light.



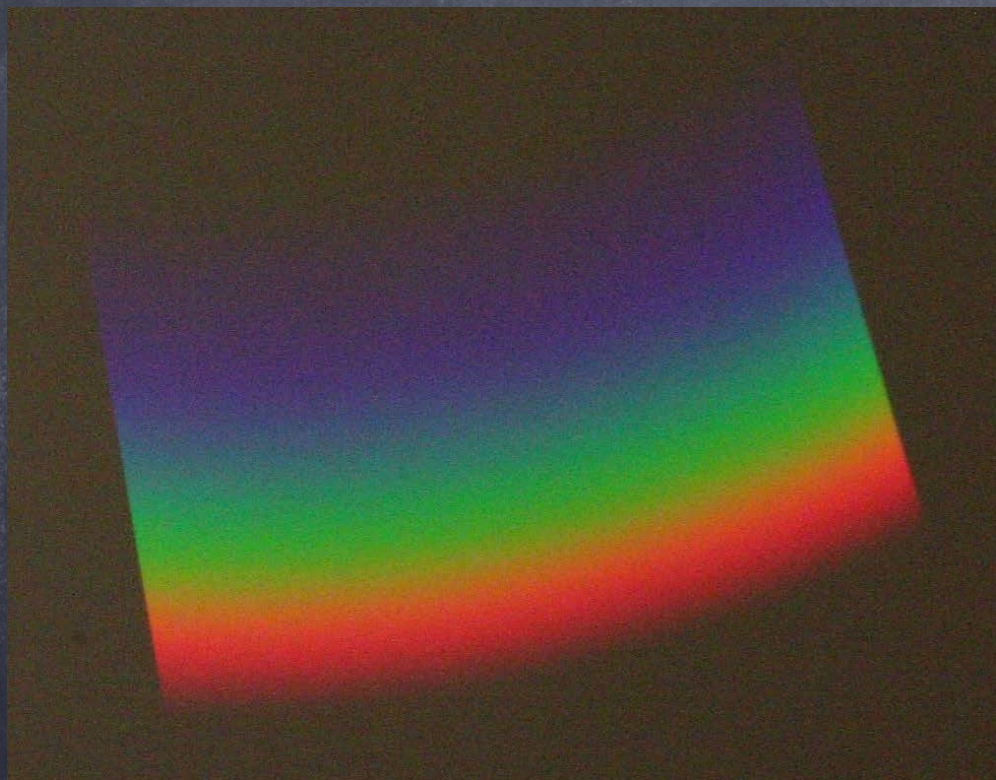
$n_1 \approx 1$ for air
 $n_2 > n_1$ in glass

low wavelength light refracts more

Experiments done
its wavelengths

where we split light into

with a prism

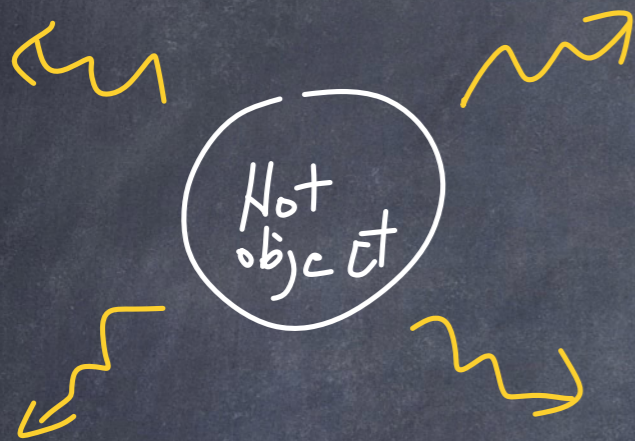


with a diffraction grating



In PHY 117, you learned that heat can be transferred by conduction, convection, & radiation (today)

Hot objects radiate EM radiation.



$$P = e\sigma AT^4$$

P : power [watts = W]

e : emissivity

σ : Stefan's Constant

$$\sigma = 5.6703 \times 10^{-8} \frac{W}{m^2}$$

T : temperature of object in K

e : 0: highly reflective
1: highly absorbant

material	e
gold	0.03
water	0.95
white paint	0.88-0.92
black paint	0.9-0.98

(emissivity also depends on shininess of material)

Object will emit radiation and absorb radiation from its surroundings.



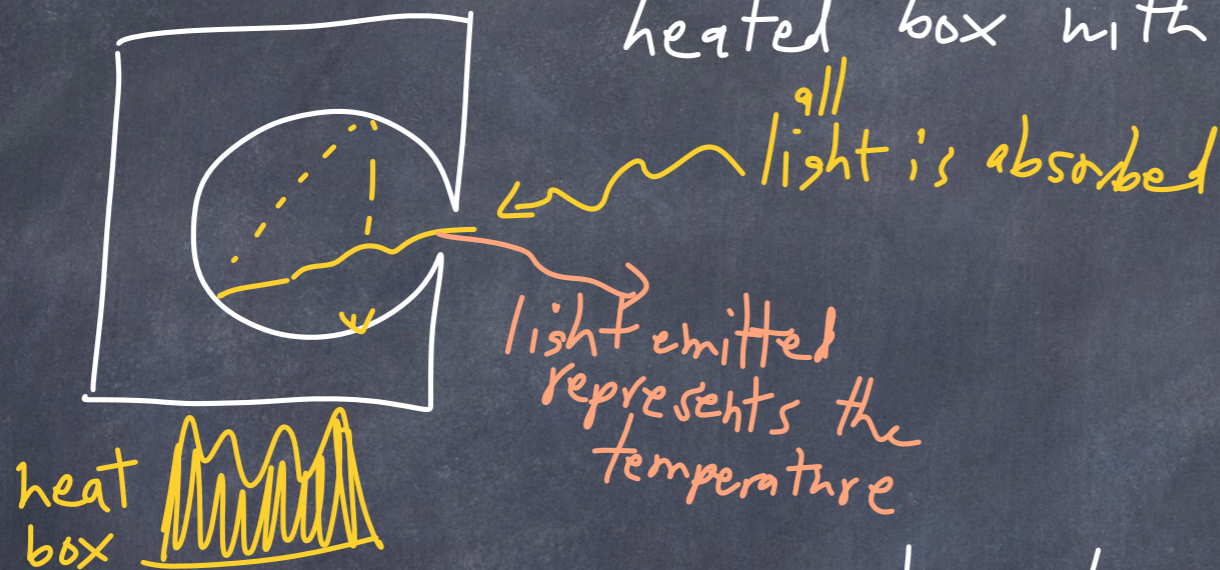
$$P = e\sigma A \left(\overset{\substack{\uparrow \\ \text{emitted}}}{T^4} - \overset{\substack{\downarrow \\ \text{absorbed}}}{T_0^4} \right)$$

IF $T > T_0$, then object will cool down with power P

IF $T_0 > T$, then object will warm up with P

If $e = 1$, object is called a perfect blackbody.
 It absorbs all radiation that it receives, it also radiates perfectly.

A perfect black body is imagined like this:
 heated box with a hole in it.

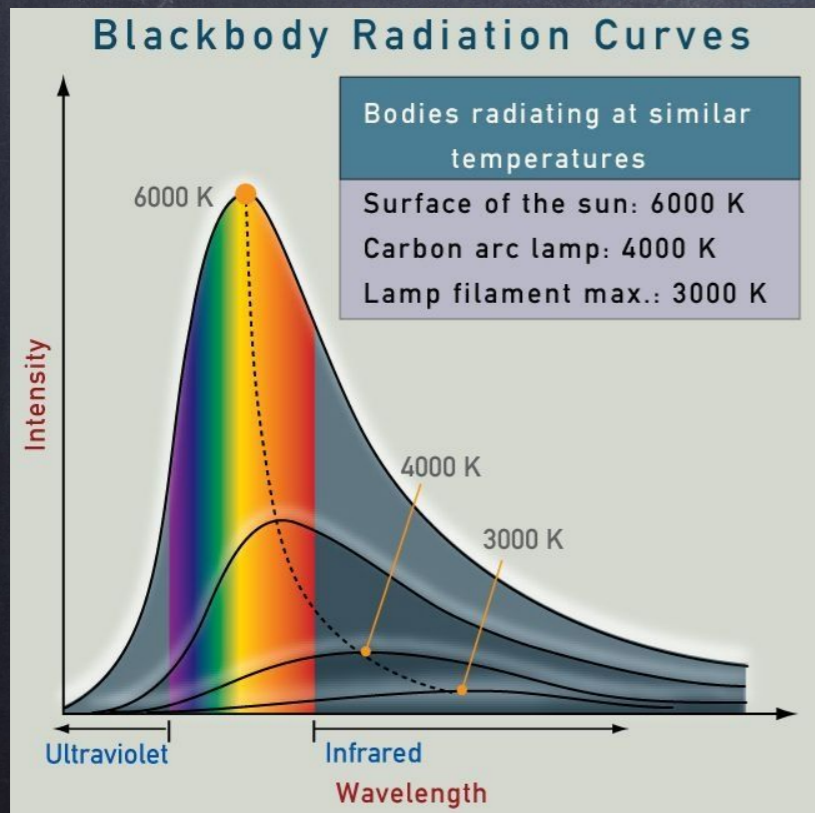


the characteristic radiation of a black body.

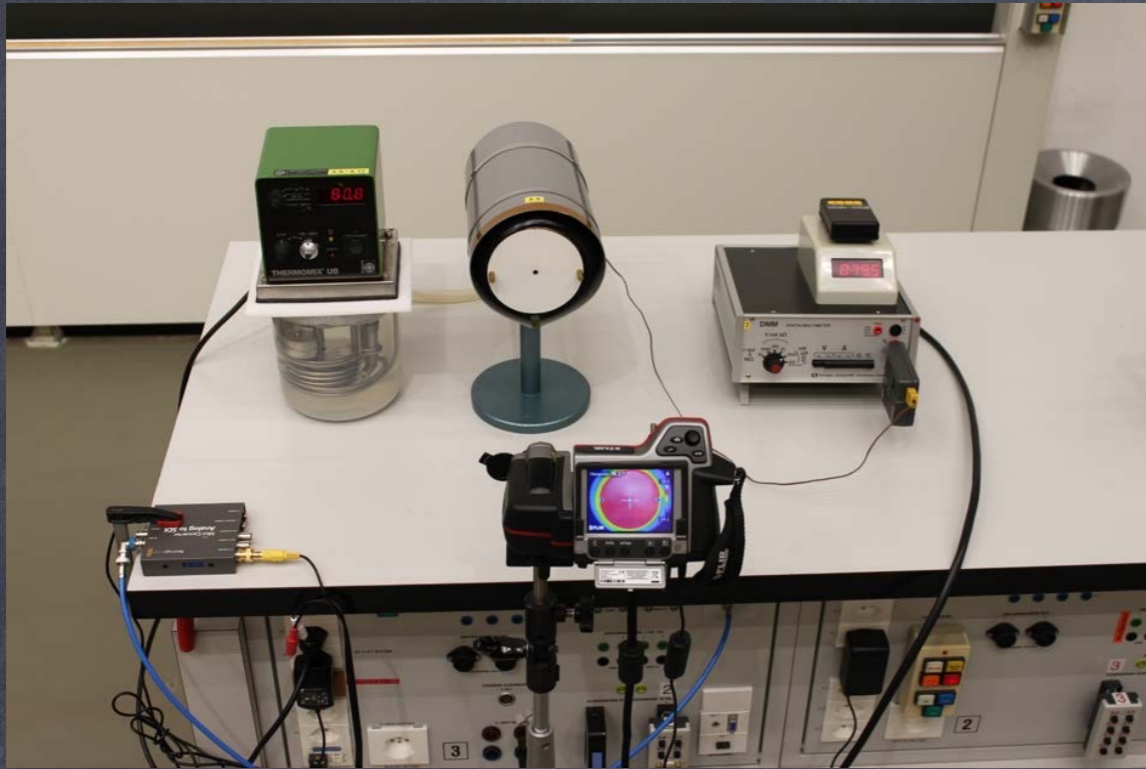
The peak wavelength depends on the temperature of the object.

$$\lambda_{\max} = \frac{2.898 \text{ mm} \cdot \text{K}}{T}$$

T : temperature in K
 Hotter temperatures emit lower wavelength light.

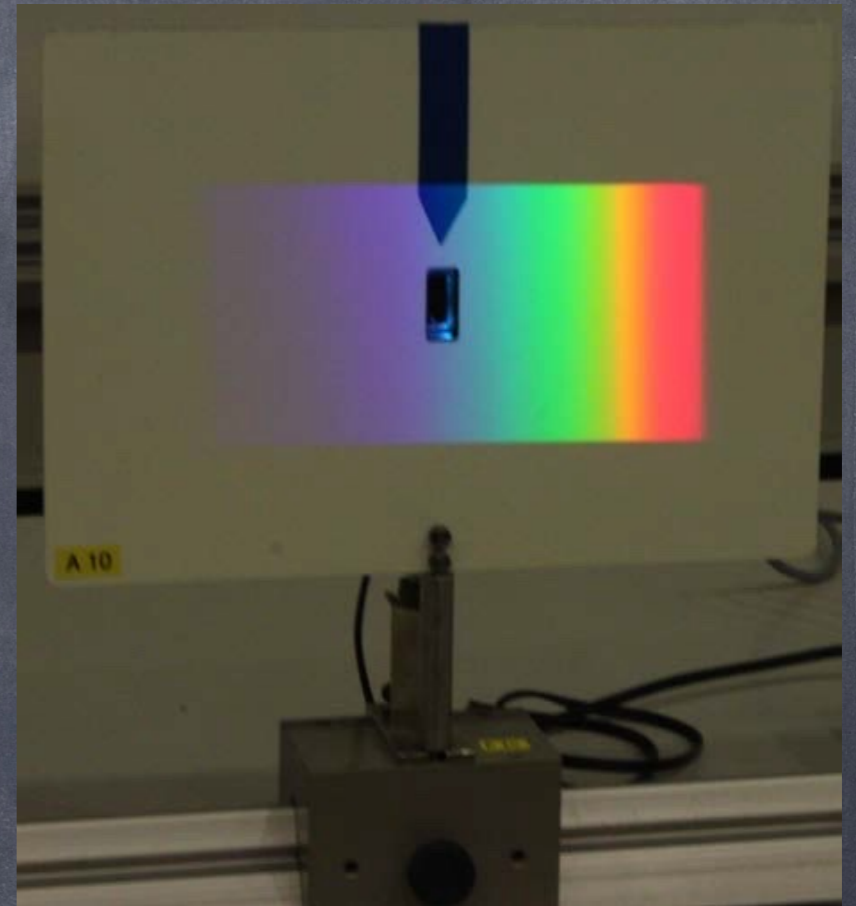


Experiments where we view blackbody radiation

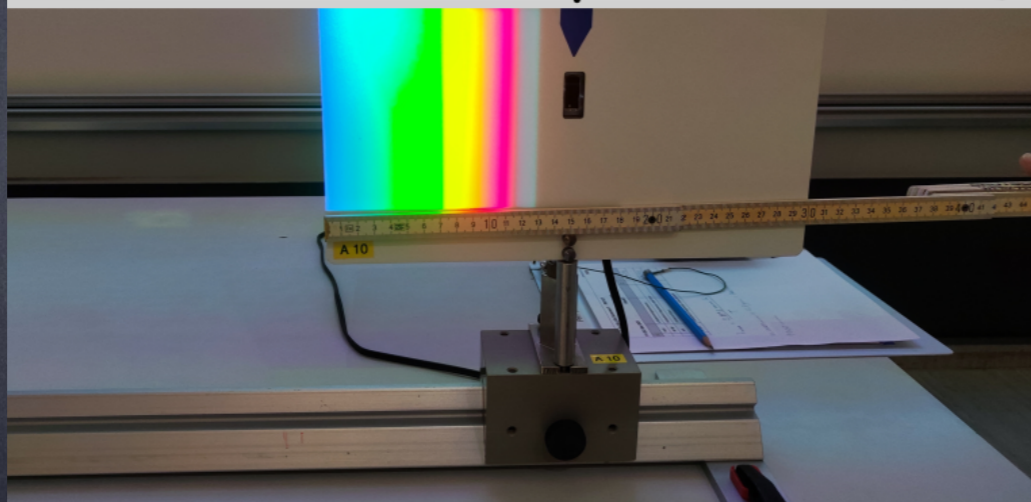
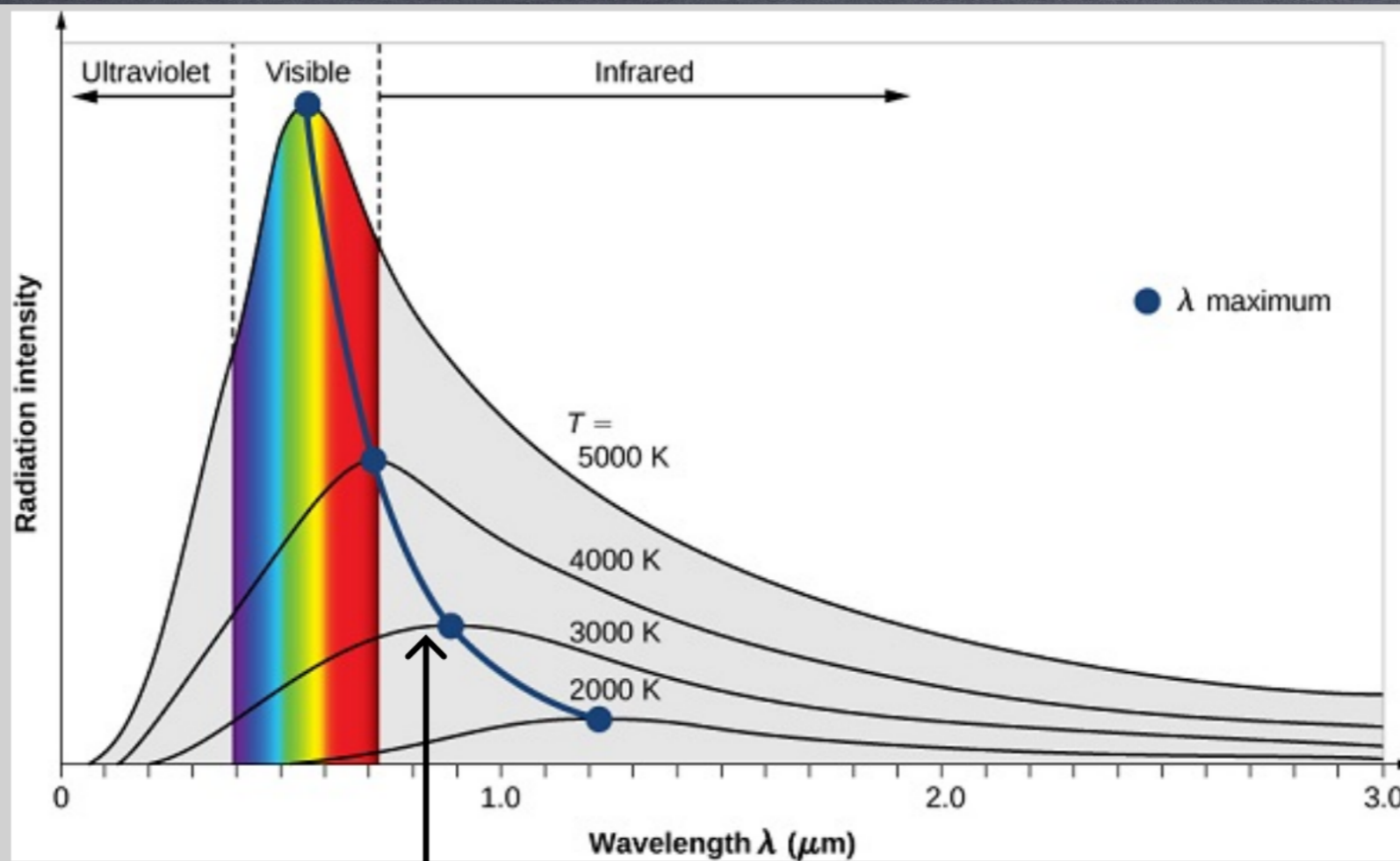


Here we see the temperature of a heated canister emitting radiation. Wien's Law lets us convert from λ_{max} to temperature assuming the value of emissivity.

Note: if we point at a material with low emissivity, the camera will mistakenly think the object is cooler than it is.



Here we measure the intensity of light emitted from a carbon arc lamp. We split the light in a prism so we can measure intensity vs. λ .



$$\frac{1000 \mu\text{m}}{1 \text{ nm}}$$

measured:

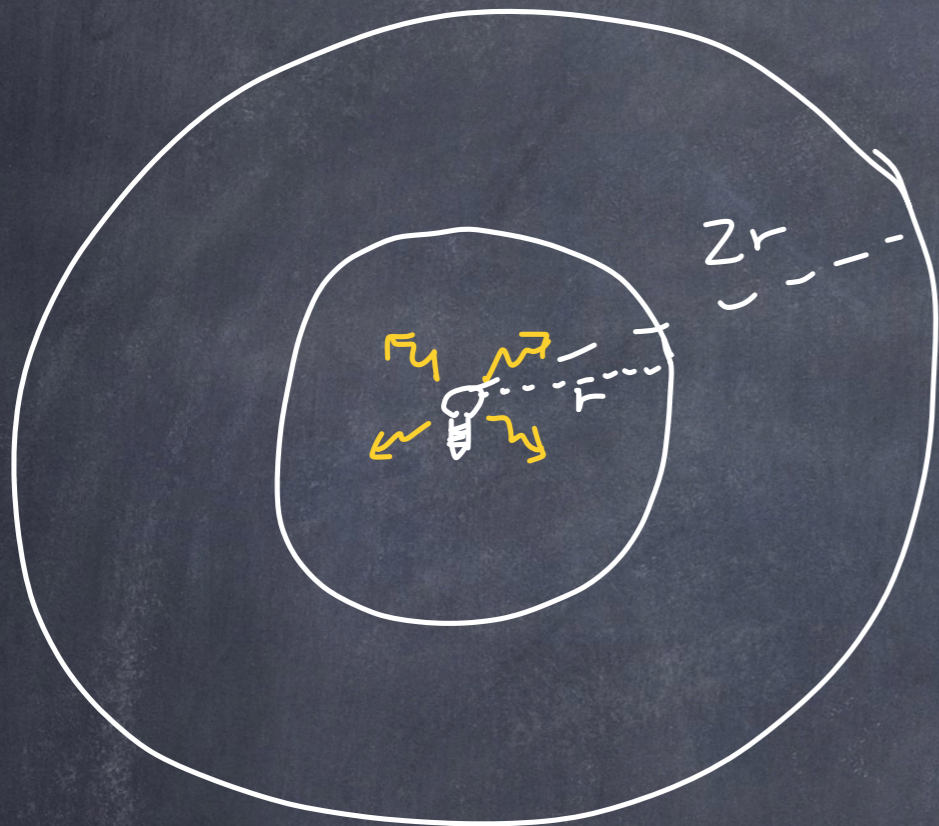
$$\lambda_{\text{max}} = 0.83 \mu\text{m} \implies T = \frac{2.898 \text{ mm} \cdot \text{K}}{\lambda_{\text{max}}} = 3500 \text{ K}$$

↑
This is actually correct according to manufacturer

Light intensity changes with distance,

$$I = \frac{\text{Power}}{\text{area}}$$

Sphere has a surface area of $4\pi r^2$



If light has an intensity of 100 W/m^2 at a distance r , what is the intensity at $2r$?

$$I_r = \frac{\text{power}}{\text{area}} = 100 \frac{\text{W}}{\text{m}^2}$$

$$\text{The total power} = (I_r)(\text{area}) = I_r (4\pi r^2)$$

$$\text{At } 2r, I = \frac{\text{power}}{\text{area}} = \frac{(100 \frac{\text{W}}{\text{m}^2})(4\pi r^2)}{4\pi (2r)^2} = \frac{100 \frac{\text{W}}{\text{m}^2}}{4}$$

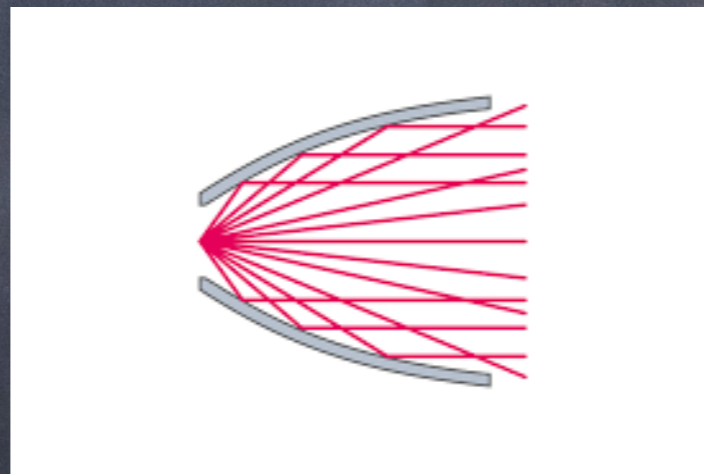
$$I = 25 \frac{\text{W}}{\text{m}^2}$$

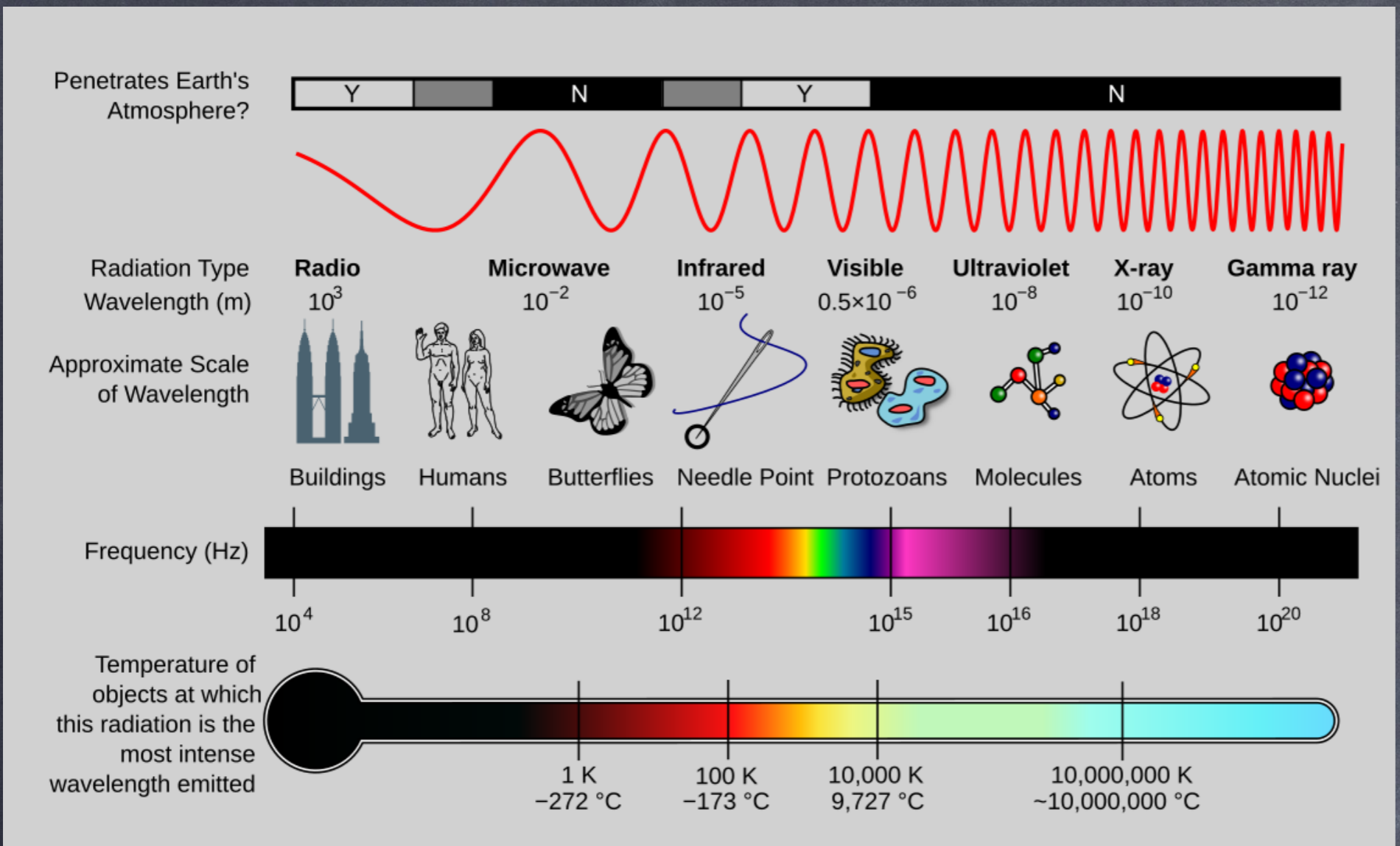
Note: efficiency of a light bulb:
 incandescent: 10-20%
 LED: 40-50%

This is the efficiency to convert electrical energy to light energy

So $(\text{power of light}) = (\text{efficiency}) * (\text{power of electricity})$

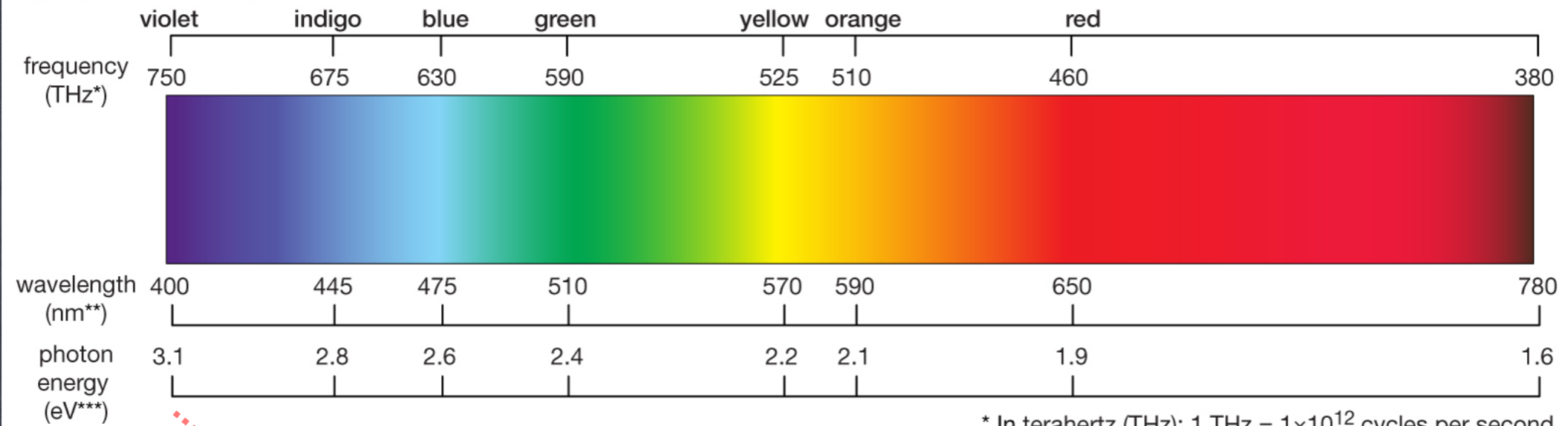
Also note that some light bulbs don't emit spherically since it is focused in one direction:





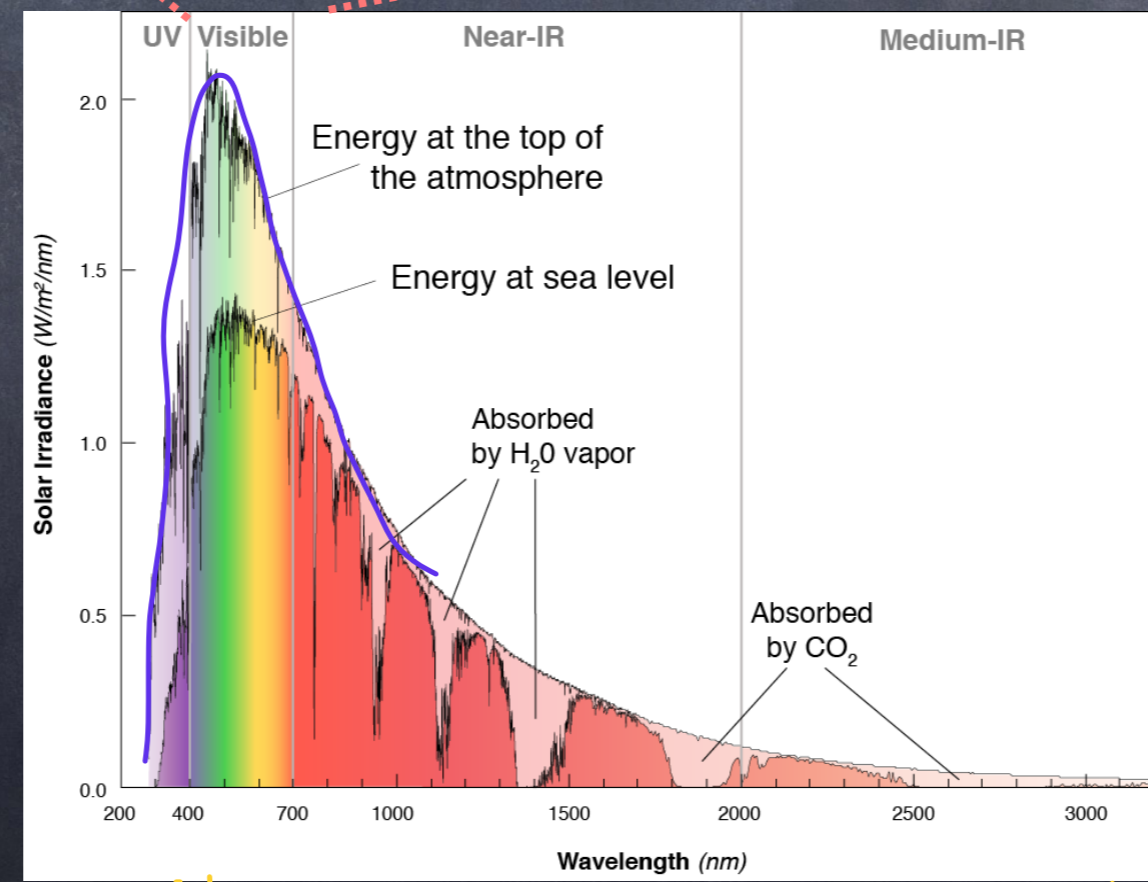
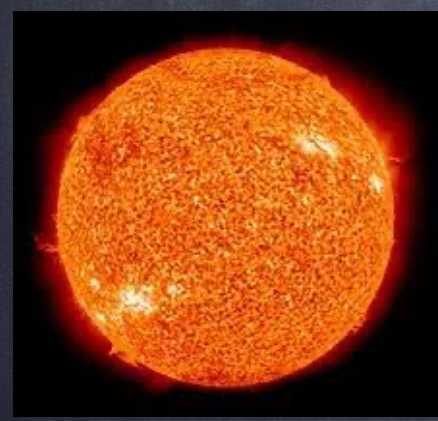
T of outer space is 2.7 K → microwave radiation
 we observe the "cosmic microwave background"
 temperature using microwave antennas

Light, the visible spectrum



* In terahertz (THz); 1 THz = 1×10^{12} cycles per second.
 ** In nanometres (nm); 1 nm = 1×10^{-9} metre.
 *** In electron volts (eV).

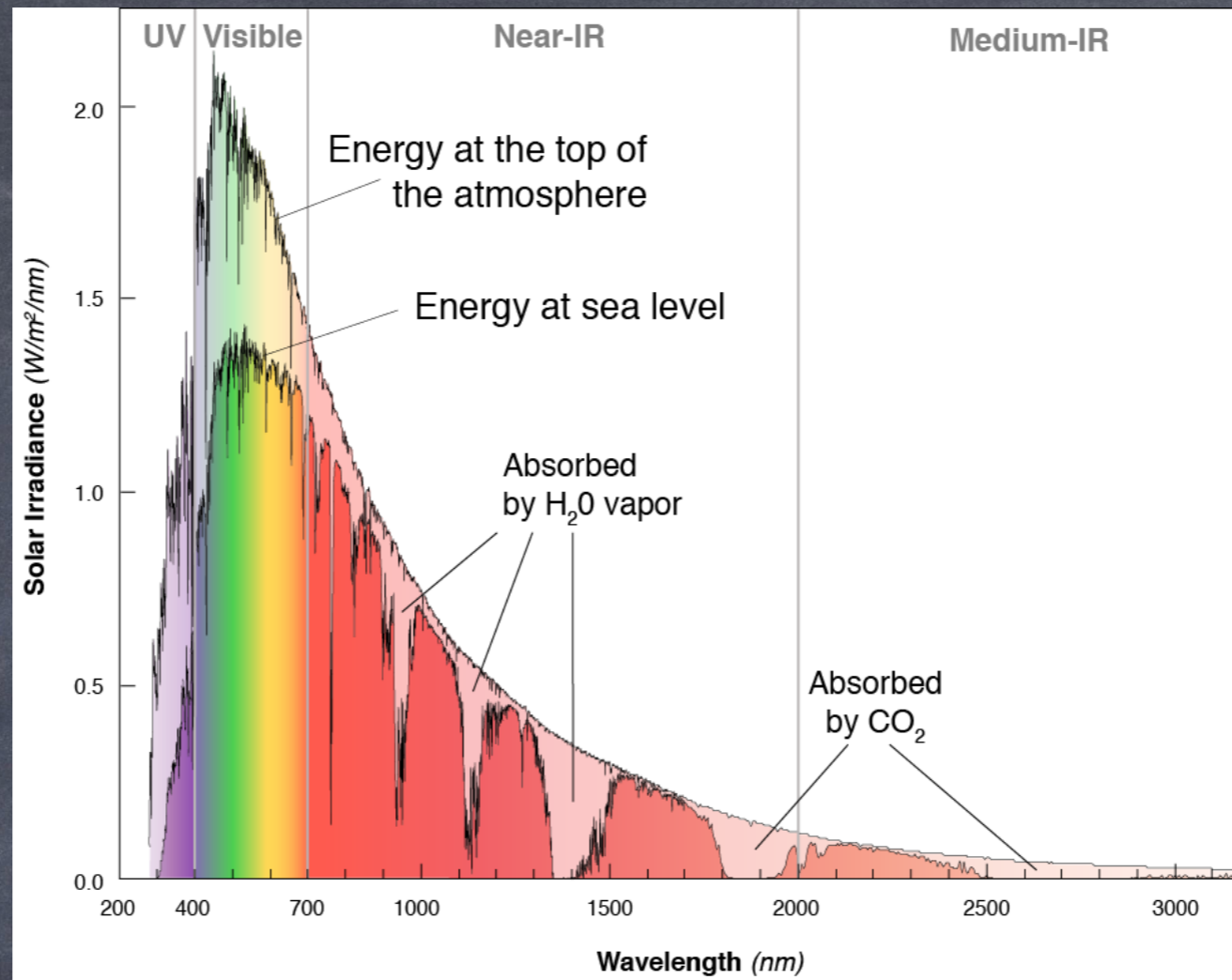
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λ_{max} for the sun ~ 500 nm

The sun radiates as a perfect blackbody

intensity vs. wavelength of sunlight

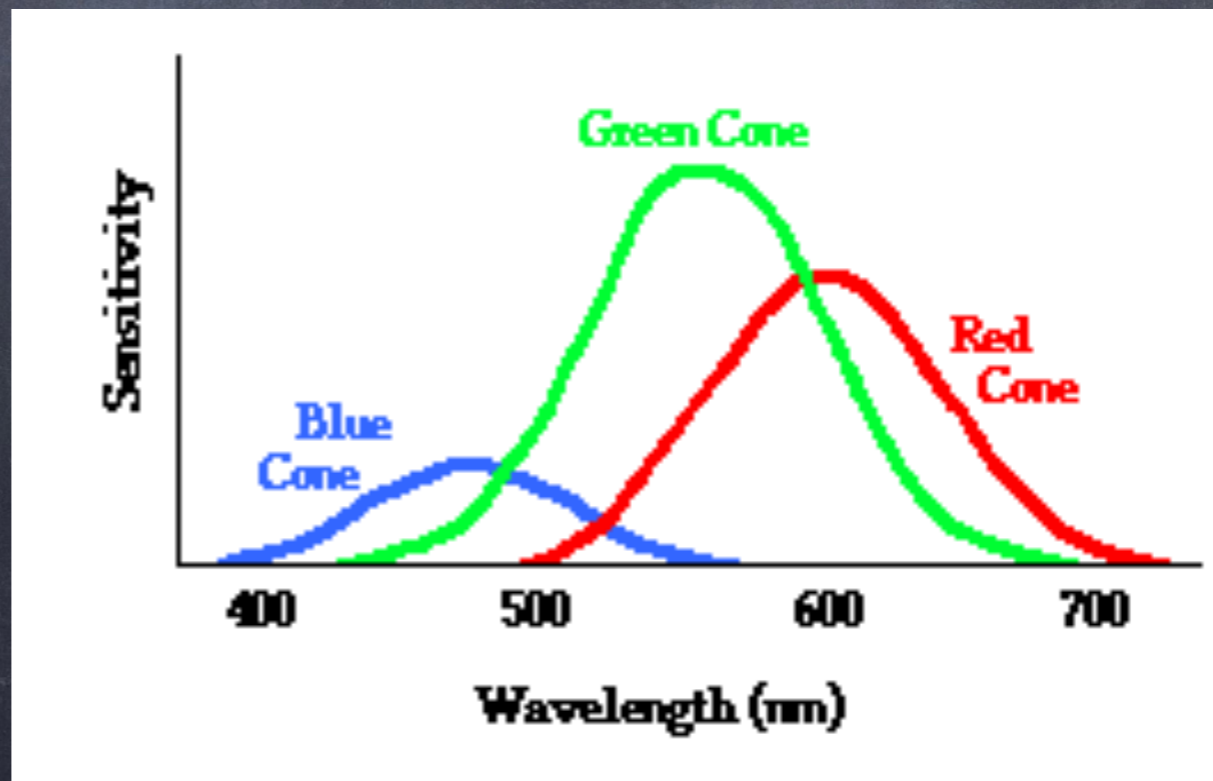


Human retina
contains rods + cones.

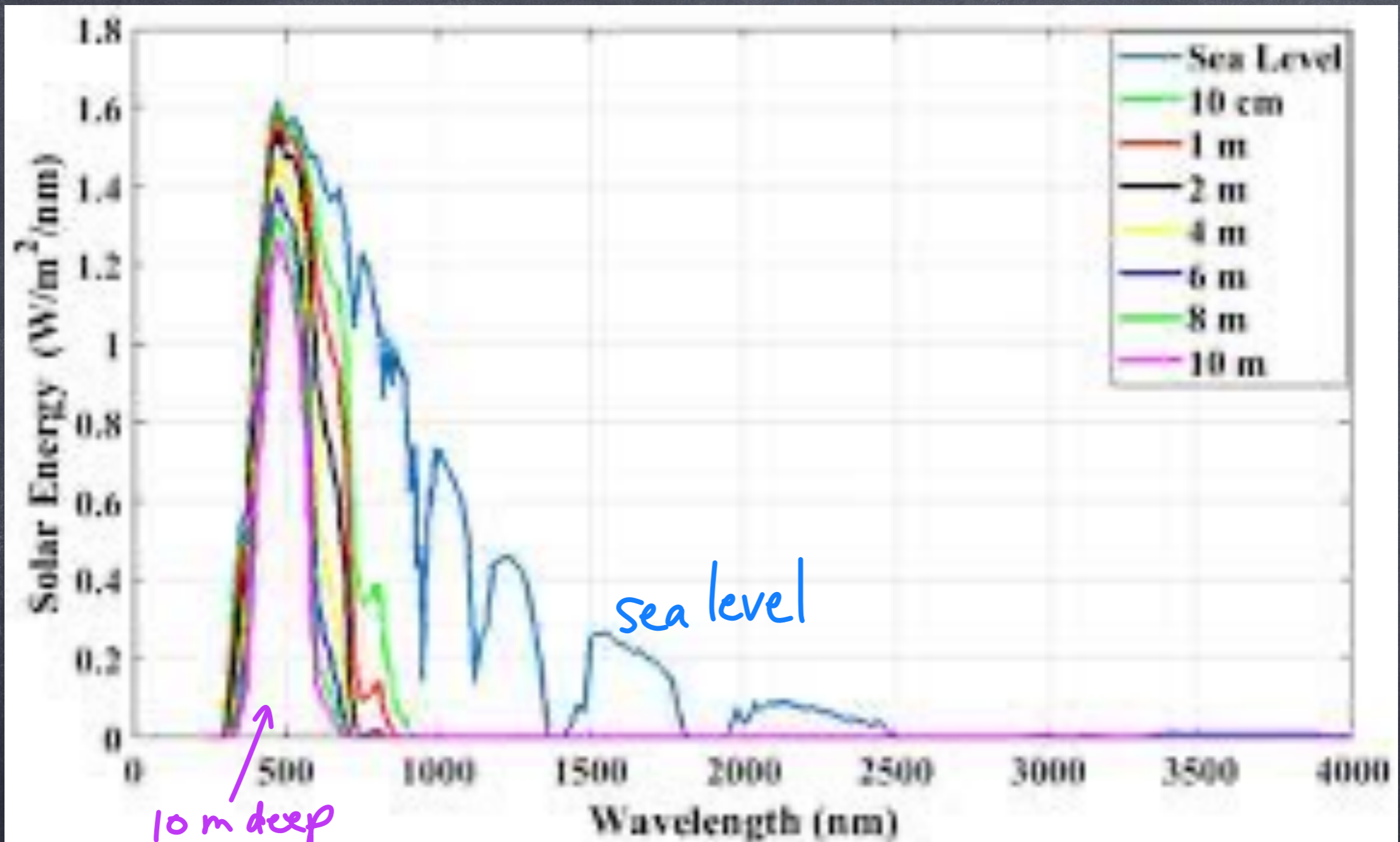
→ Rods measure the
intensity of light.

→ cones measure light
color.

Human eye has 3 cones;
sensitive to different λ of light.



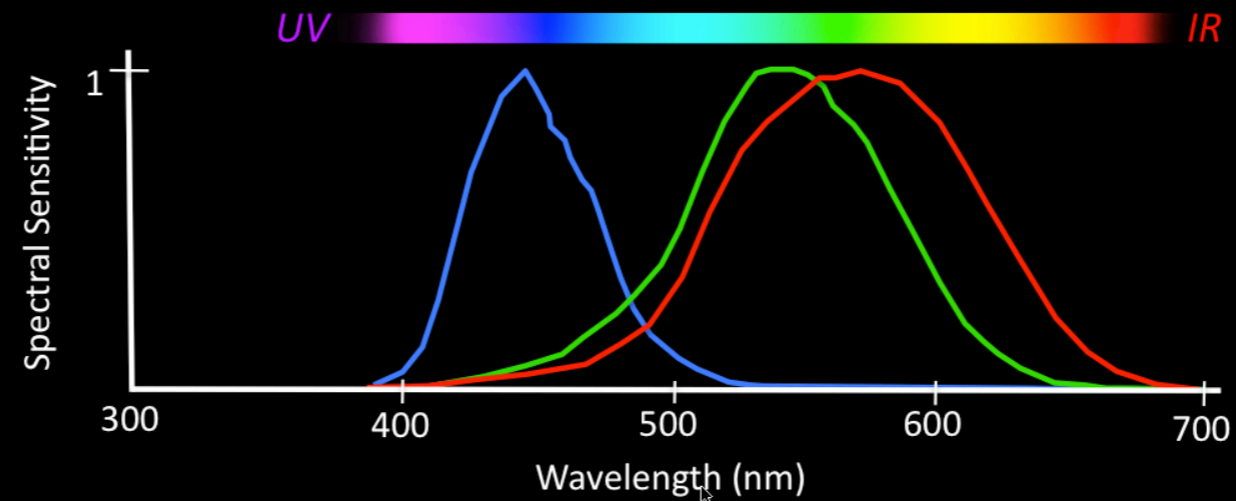
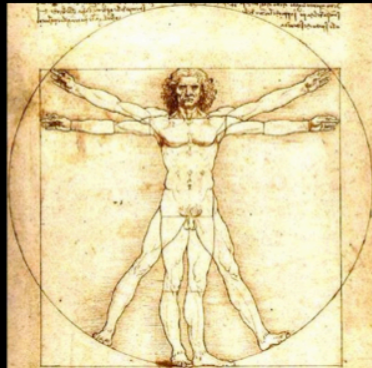
Intensity of sunlight below sea level.



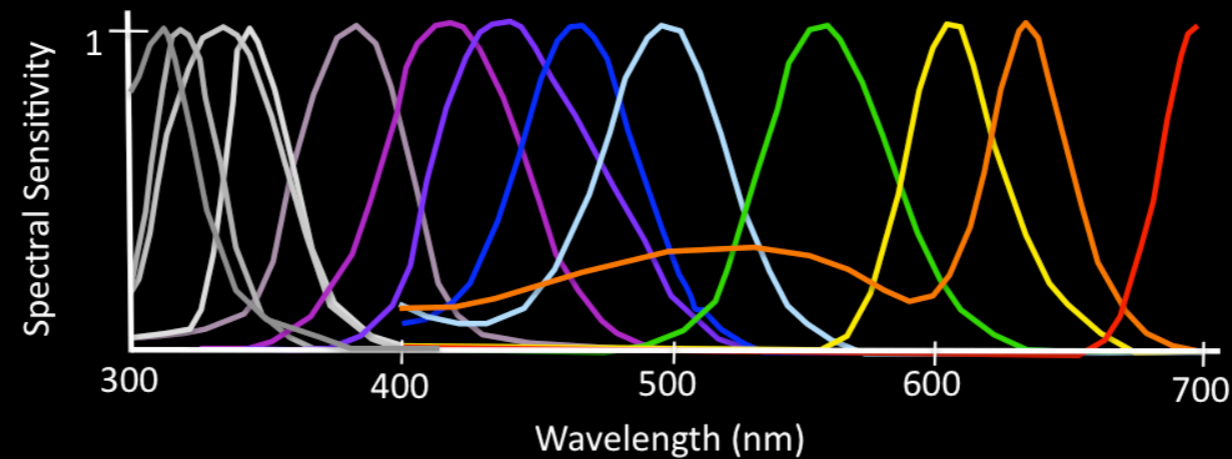
higher wavelength light is reflected or absorbed.

Mantis Shrimp: Extraordinary Eyes

Homo sapiens



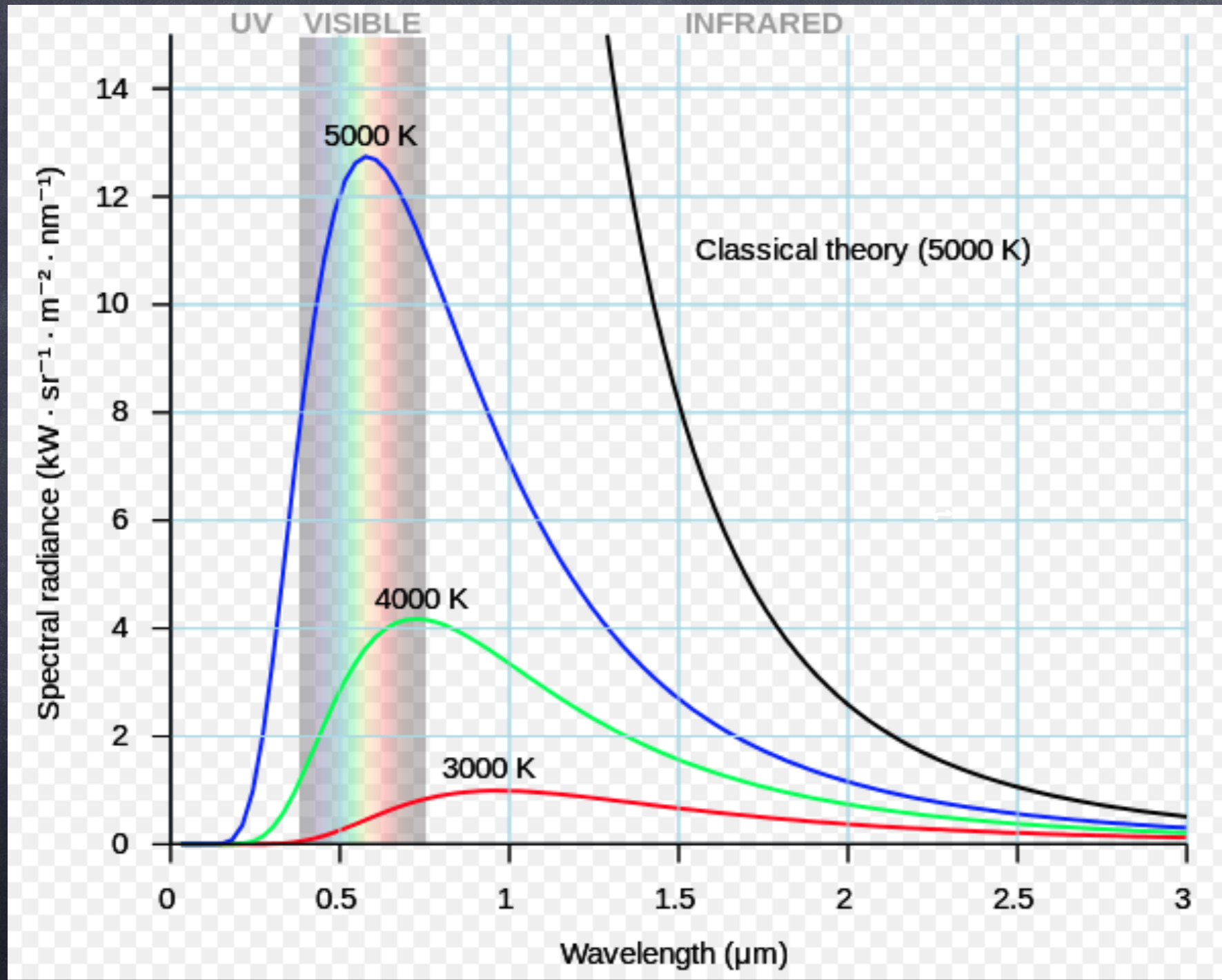
Neogonodactylus oestedii



Marshall *et al.*, 2007; Marshall and Oberwinkler, 1999

extra [↑]sensitivity to low wavelengths

The mystery of blackbody radiation. (19th century)



classical theory predicts an infinite amount of low wavelength light from a blackbody radiation.
(Not what is observed.)

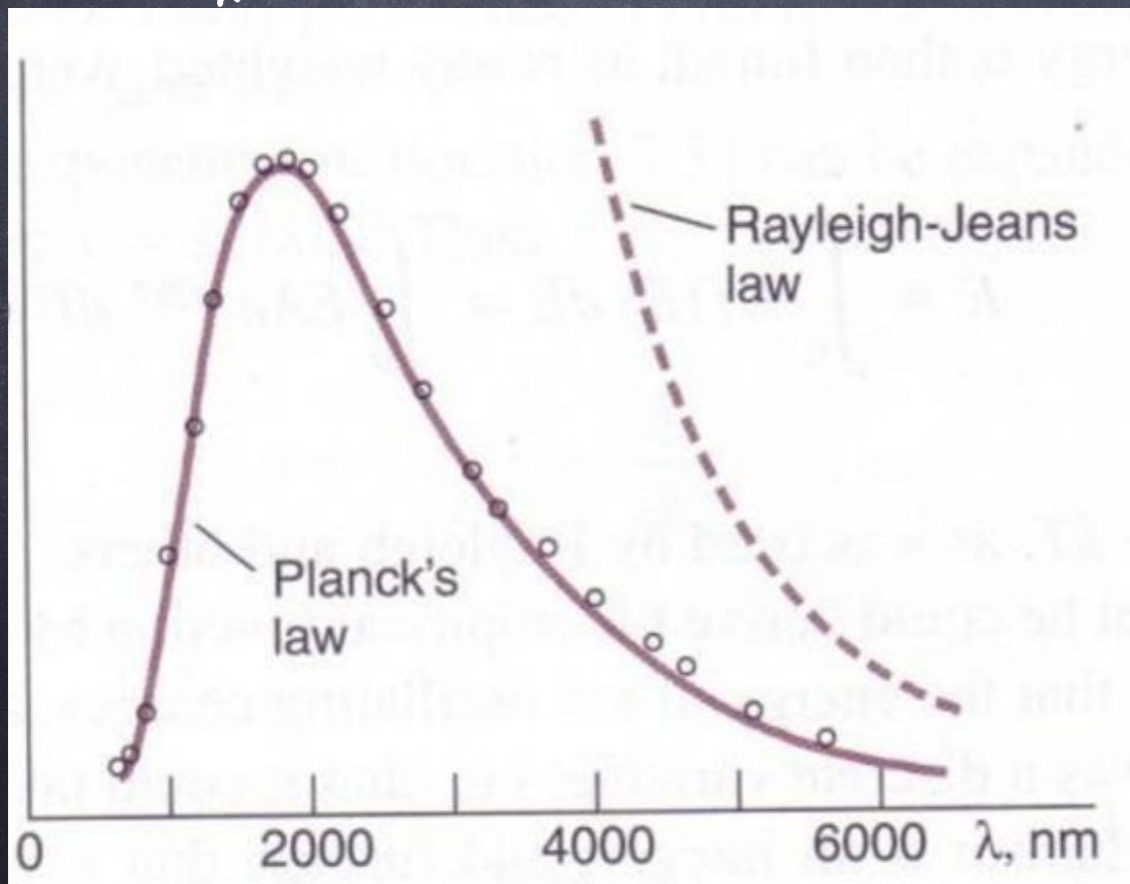
This was solved by Planck.

$$\text{Intensity} = I = \frac{2\pi c^2 h}{\lambda^5} \frac{1}{e^{hc/\lambda kT} - 1}$$

k: Boltzmann constant $k = 1.38 \times 10^{-23} \text{ J/K}$
 same as $(PV = n k T)$

h: Planck constant = $h = 6.261 \times 10^{-34} \text{ J}\cdot\text{s}$

Planck's
 ↓

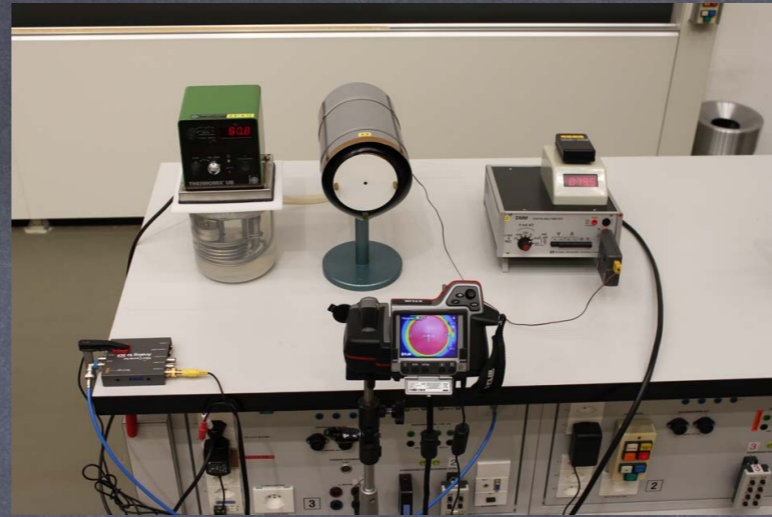


← classical theory
 solution of Planck:
 Considers that a blackbody radiates light as if little harmonic oscillators, Each one with energy $E = \frac{hc}{\lambda}$

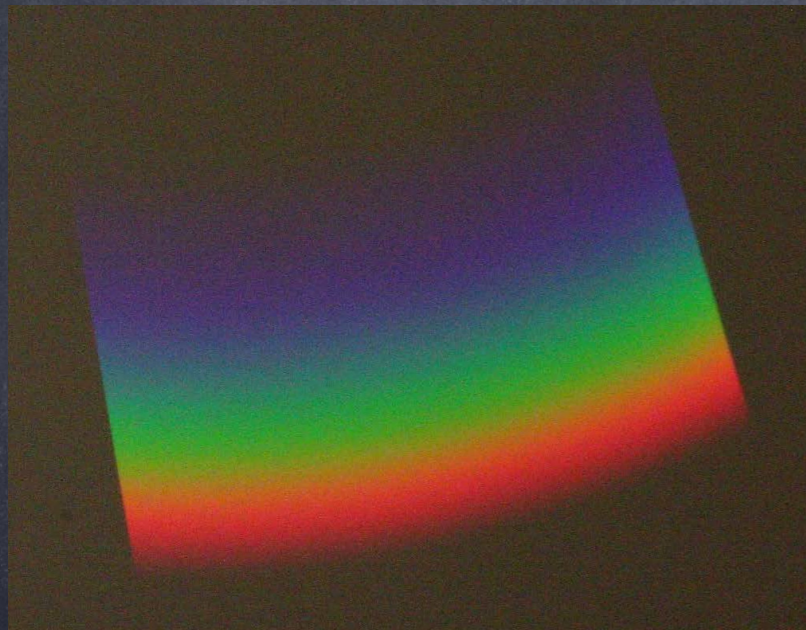
This worked, but no one understood why.



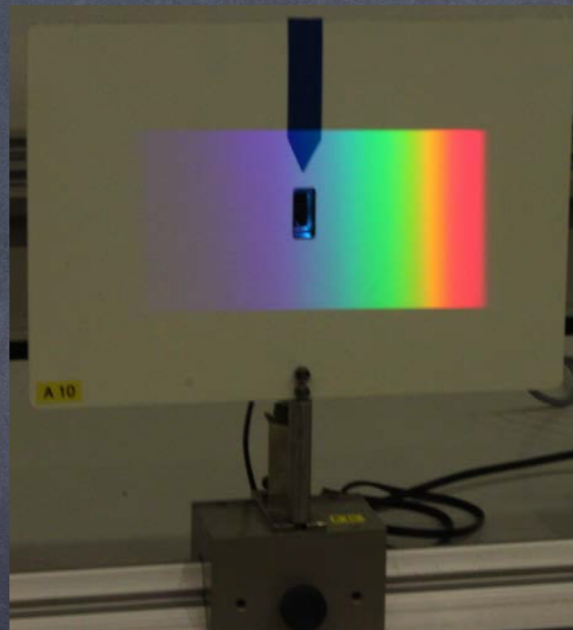
W51



A8



W100



A10



W101